

# PHYS 431/531: Galactic Dynamics

## Fall 2011, Homework #2

(Due October 20, 2011)

1. (a) Given the definitions of the Oort constants  $A$  and  $B$  presented in class (Eqs. 2.13 and 2.16 in the text),

$$A = -\frac{1}{2}R \left(\frac{V}{R}\right)' \quad B = -\frac{1}{2} \frac{(RV)'}{R},$$

verify that  $A + B = -V'(R_0)$  and  $A - B = V_0/R_0$ , where  $V(R)$  is the Galactic rotation law,  $R_0$  is the distance from the Sun to the Galactic center, and  $V_0 = V(R_0)$ .

- (b) Consider the spherically symmetric density distribution  $\rho$  given by

$$\rho(R) = \rho_0 \left(1 + \frac{R^2}{a^2}\right)^{-1}.$$

Derive an expression for the mass inside radius  $R$ . What is the circular orbital speed at radius  $R$ ? Hence determine the form of  $A(R)$  and  $B(R)$  for  $R \gg a$ .

2. Air at sea level has density  $\rho = 1.2 \text{ kg/m}^3$  and sound speed  $v_s = 330 \text{ m/s}$ .
- (a) What is its Jeans length? What is the Jeans mass?
- (b) By how much does the self-gravity of air change the frequency of a sound wave of wavelength 1 m?
- (c) By what fraction (relative to the “short-wavelength” result  $f = v_s/\lambda$ ) does the self-gravity of an interstellar cloud change the frequency of a sound wave with wavelength equal to half the Jeans length?
3. (a) A particle is dropped (from radius  $a$  with zero velocity) into the gravitational potential corresponding to a static homogeneous sphere of radius  $a$  and density  $\rho$ . Calculate how long the particle takes to reach the center of the sphere. [Hint: the equation of motion is  $d^2r/dt^2 = -GM(r)/r^2$ .]
- (b) Calculate the time required for a homogeneous sphere of radius  $a$  and density  $\rho$  with no internal pressure support to collapse under its own gravity. [Apply the previous equation of motion to a particle on the surface.]
4. Interstellar gas in many galaxies is in virial equilibrium with the stars, in that the rms speed of the gas particles is the same as the rms stellar speed. Consider a large elliptical galaxy with a virial radius of 100 kpc and a mass of  $10^{12}$  solar masses. Calculate the rms stellar velocity using the virial theorem. Hence estimate the temperature of the interstellar gas, assuming that it is composed entirely of hydrogen.